# Notes on the Riemann hypothesis and abundant numbers

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<code>RH\_abundant.tex</code> ('NOTES' OPTION) TYPESET IN <code>PDFIATEX</code> ON A LINUX SYSTEM

# Introduction

★ In [1], Lagarias showed the equivalence of the Riemann hypothesis (RH) to a condition on harmonic sums, namely

 $\mathsf{RH} \Leftrightarrow \sigma(n) \leqslant H_n + \exp(H_n) \log(H_n) \ \forall n$ 

 $\star$  Robin [3] had already shown that

 $\mathsf{RH} \Leftrightarrow \sigma(n)/n < e^{\gamma} \log \log (n) \quad \text{for} \quad n > 5040$ 

which is probably more convenient for numerical tests  $\star$  so to disprove RH, we need to find n with large  $\sigma(n)/n$ 

# Notation

- ★ n is always a positive integer; p is always a prime ★  $p_i = i$ th prime  $(p_1=2, p_2=3,...)$
- $\star$  harmonic sum for  $n \ge 1$  is  $H_n = \sum_{i=1}^n i^{-1}$
- $\star$  sum of divisors is  $\sigma(n) = \prod_i \frac{p_i^{a_i+1}-1}{p_i-1}$  for  $n = \prod_i p_i^{a_i}$ ,  $(a_i > 0)$

$$\star \sigma(n)/n = \prod_i \frac{p_i - p_i^{-a_i}}{p_i - 1}$$

- **\star** I define  $\rho(n) \equiv \sigma(n)/n$
- $\star~e^{\gamma}=1.78107241799019798523650410310717954\ldots$  , where  $\gamma~$  is Euler's constant
- $\star$  we will deal with n too large to represent in the computer ( $\approx 10^{(10^{10})}$ ), but luckily we have its prime factorization which suffices

# Superabundant numbers [2]

- $\star~n$  is superabundant (SA) if  $\sigma(k)/k < \sigma(n)/n$  for all k < n
- ★ if  $n = 2^{a_2} 3^{a_3} \dots m^{a_m}$  is SA, where m is the maximal prime factor, then  $a_2 \ge a_3 \ge \dots \ge a_m$
- $\star$  if  $1 < j < i \leq m$ , then  $|a_i \lfloor a_j \log_i j \rfloor| \leq 1$
- $\star$   $a_m = 1$  unless n = 4 or 36
- $\star i^{a_i} < 2^{a_2+1}, \ i \geqslant 2$
- $\star \ m \sim \log(n)$
- \* SA numbers, with CA numbers in red:  $2, 2^2, 2.3, 2^2.3, 2^3.3, 2^2.3^2, 2^4.3, 2^2.3.5 \dots$

# The work of Robin [3]

★ Theorem:

 $\mathsf{RH} \Leftrightarrow \rho(n) < e^{\gamma} \log \log (n) \quad \text{for} \quad n > 5040$ 

**★** Theorem: independently of RH, except for n = 1, 2, 12:

$$\rho(n) < e^{\gamma} \log \log \left(n\right) + \frac{\left[7/3 - e^{\gamma} \log \log \left(12\right)\right] \log \log \left(12\right)}{\log \log \left(n\right)}$$

The numerator in the last term is about 0.6482

\* But note  $\limsup_{n\to\infty} \frac{\sigma(n)}{n\log\log(n)} = e^{\gamma}$ 

#### The structure of the set of CA numbers

- ★ Defn: *n* is colossally abundant (CA) if there exists  $\epsilon > 0$  such that  $\sigma(n)/n^{1+\epsilon} \ge \sigma(k)/k^{1+\epsilon} \quad \forall k > 1$
- ★ Robin [3] has given the following precise description, due originally to Erdős & Nicolas [6]
- $\star$  we first form the set E of critical  $\epsilon$  values:

$$E_p \equiv \bigcup_{\alpha=1,2,3,\dots} \{ \log_p \left( 1 + \frac{1}{\sum_{i=1}^{\alpha} p^i} \right) \}$$
$$E \equiv \bigcup_p E_p$$

\* we label the elements of E in decreasing order:  $\epsilon_1 = \log_2(3/2) > \epsilon_2 = \log_3(4/3) > \epsilon_3 = \log_2(7/6) > \dots$ 

#### The structure of the set of CA numbers cotd.

(a) if  $\epsilon \notin E$ ,  $\sigma(n)/n^{1+\epsilon}$  has a unique maximum attained at the number  $n_{\epsilon}$  with prime exponents given by

$$a_p(\epsilon) = \left\lfloor \log_p \left( \frac{p^{1+\epsilon} - 1}{p^{\epsilon} - 1} \right) \right\rfloor - 1$$

- (b) if  $\epsilon$  satisfies  $\epsilon_{i+1} < \epsilon < \epsilon_i$  for i = 1, 2, 3, ..., then  $n_{\epsilon}$  is constant and we call it  $N_i$ . We have  $N_1 = 2, N_2 = 6, ...$
- (c) if the sets  $E_p$  are pairwise disjoint (which is likely, but not certainly known), then the set of CA numbers is equal to the set of  $N_i, i = 1, 2, 3, \ldots$ . If this is the case,  $\sigma(n)/n^{1+\epsilon}$  attains its maximum at the two points  $N_i$  and  $N_{i+1}$
- (d) if the sets  $E_p$  are not pairwise disjoint, then for each  $\epsilon_i \in E_q \cap E_p$ ,  $\sigma(n)/n^{1+\epsilon_i}$  attains its maximum at the four points  $N_i$ ,  $qN_i$ ,  $rN_i$  and  $N_{i+1} = qrN_i$

# Alaoglu & Erdős [2]

- $\star$  these authors have a slightly stronger definition:
- $\star$  Defn: *n* is colossally abundant if

$$\begin{array}{lll} \sigma(n)/n^{1+\epsilon} & > & \sigma(k)/k^{1+\epsilon} & \mbox{for} & 1 \leqslant k < n \\ \sigma(n)/n^{1+\epsilon} & \geqslant & \sigma(k)/k^{1+\epsilon} & \mbox{for} & k > n \end{array}$$

- ★ the effect of this is to make a unique choice from the 2 or 4 possibilities in cases (c) and (d) above. But I will perform my computations with the Robin definition
- ★ we will call these numbers strongly colossally abundant (SCA) if it is necessary to distinguish them from ordinary CA numbers

# Position of critical $\epsilon$ values



critical  $\varepsilon$  values

The vertical lines mark the critical  $\epsilon$  values arising from the small primes on the y axis

# Density of critical $\epsilon$ values



#### Dependence of maximal prime on n



# Method

- ★ it is known that if RH is false, there will be a violation of Robin's inequality which is a CA number . . .
- $\star$  for a range of small  $\epsilon>0,$  I compute n by the A&E formula
- ★ I compute RHS-LHS of Robin's inequality (let's call this the *deviation*  $\delta(n) \equiv e^{\gamma} \log \log (n) \rho(n)$ ), and look for any violations (i.e.  $\delta < 0$ )
- \* we can also plot  $\eta(n) \equiv e^{\gamma} \rho(n) / \log \log (n)$
- $\star$  the following plots show the behaviour observed so far (to about  $\epsilon = \exp(-25)$ )

# Computational difficulties

- ★ the exponents  $a_p(\epsilon)$  must be computed in interval arithmetic, to ensure they are correct and not corrupted by roundoff error. This means not just high precision arithmetic, but dynamically varying precision
- $\star$  millions of primes are needed. Typically the *n* we deal with have a huge tail of many primes to the power 1. It is fastest to precompute primes with a sieve, but then much storage is required.
- $\star$  how do we vary  $\epsilon$  to not miss any SCA numbers? (DONE)
- $\star$  how to we compute explicit examples of WCA numbers?
- $\star$  there are many other difficulties . . .

#### Computational strategy

We keep a list z of records, containing: a prime p,  $\log p$ , its exponent a, and a critical  $\epsilon_c$ , which is the value of  $\epsilon$  at which this exponent will next change (as  $\epsilon$  is decreased). We exclude 1 exponents, which are counted by *ones*. We first initialize:

▷ fix  $0 < \epsilon_{start} \leq 1$ . Then, for each prime p, compute  $a = \left\lfloor \log_p \left( \frac{p^{1+\epsilon_{start}-1}}{p^{\epsilon_{start}-1}} \right) \right\rfloor - 1$ and store it in the z list if  $a \ge 2$ . If a = 1, just increment the variable ones. Stop when a = 0. During this p loop, also update  $\log(n)$  and  $\rho(n)$ 

Then each step of the main loop consists of determining which of possible events A, B, or C occurs:

- ▷ A: a new prime (with exponent 1) is added, so we increment 'ones'. This happens when  $\epsilon_{ext} = \log_p (1+p)$  is maximal, where p is the new prime
- ▷ B: the first prime with exponent 1 has its exponent raised to 2. This happens when  $\epsilon_{inc} = \log_p \left( \frac{p+1+1}{p} / \frac{p+1}{p+1} \right)$  is maximal, where p is the prime in question
- ▷ C: a prime with exponent  $\ge 2$  has its exponent incremented. This happens when  $\epsilon_{\max} = \log_p \left( (1-p^{a+1})/(p-p^{a+1}) \right)$  is maximal, where p is the prime in question and a its exponent

### Prime generation

For simple tests, just use a lookup list. With the BERNSTEIN option,  $crit_eps14.c$  uses Bernstein's quadratic sieve. Here we cannot just look up any prime; rather we have a function to get the next prime and advance the internal state (also to peek at the current prime without advancing). So the strategy is to use 2 prime generators - one (pg1) for the z list (which only grows slowly), and another (pg2) to see if the list of 1s needs extending. pg2 goes a long way, but we rely on Bernstein's code to keep it efficient.

# Questions

- $\star$  how does *n* depend on  $\epsilon$ ?
- $\star$  how does the largest prime needed depend on  $\epsilon$ ?
- $\star$  how does  $\delta$  depend on  $\epsilon$ ?
- ★ interesting observation: when p is large and  $\epsilon$  is small (the situation we are interested in), in the exponent formula of Alaoglu and Erdős it is already sufficient to use the first term of a Taylor expansion in  $\epsilon$ , namely as  $\epsilon \rightarrow 0^+$ ,

$$\log_p\left(\frac{p^{1+\epsilon}-1}{p^{\epsilon}-1}\right) = \log_p\left(\frac{p-1}{\log p}\right) - \log_p(\epsilon) + \mathcal{O}(\epsilon)$$

already has error less than 1/2, so the floor is the correct integer. How can we exploit this?

★ in the following plots the computed data is in red and hypothesized fits or trends are in blue

#### Dependence of n on $\epsilon$



This shows that  $\log \log (n)$ at SCA numbers n appears to be asymptotically a linear function of  $-\log \epsilon$ . The line -1.779+0.947xis guesswork

### Density of CA numbers



#### Dependence of $\delta$ on n



This shows that  $\log \delta$  appears to be asymptotically a linear function of  $x = \log \log (n)$ 

#### Dependence of $\delta$ on n



This is the last 100000 values of the previous plot

#### Deviation of $\log \delta$ from a best-fit line



This shows the difference between  $\log \delta$  and a conjectured best-fit line a - x/2, where  $x \equiv \log \log (n)$ 

# References

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